

OH Suppression Near IR Spectroscopy for the LBT

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LBTC Users Meeting March 23-23, 2014
Tucson Arizona

Concept Introduction

Provide a low resolution $R \sim 500$ near infrared spectrometer for faint object spectroscopy at the LBT.

Utilize the excellent LBT AO system to minimize the size of the spectrometer.

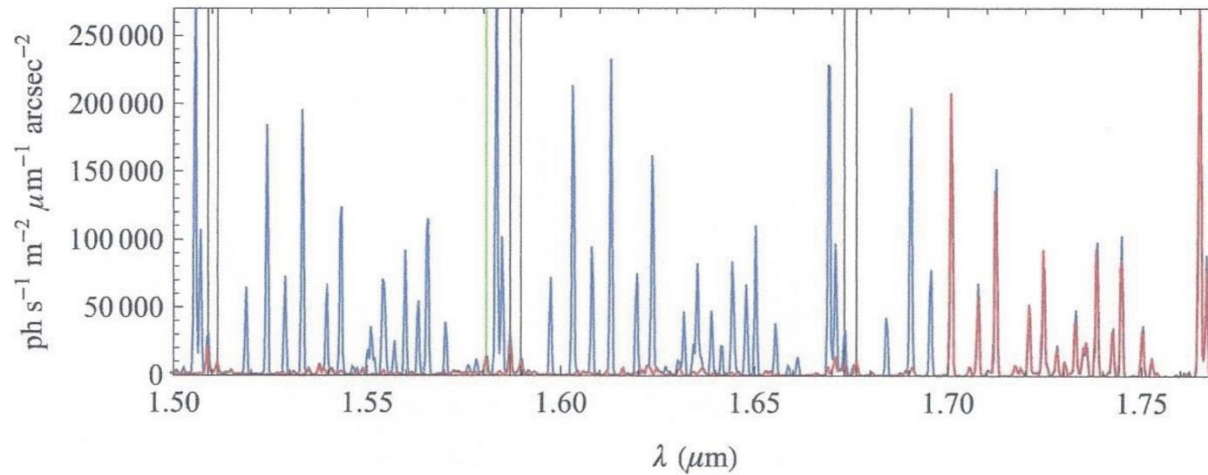
Employ Bragg Grating Fibers (BGF) to suppress the intense OH emission lines that have been the perpetual bane of low resolution near IR spectroscopy.

BGFs have been developed in Australia

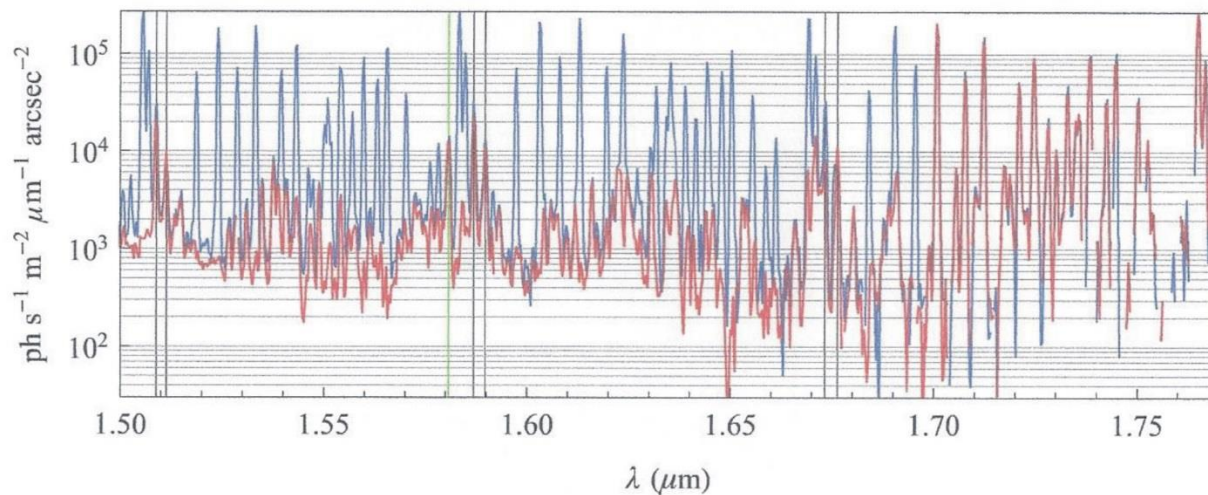
AIP is also developing the capacity to manufacture BGFs

Fibers can be placed to utilize both mirrors of the LBT.

Example of OH Suppression



(a)

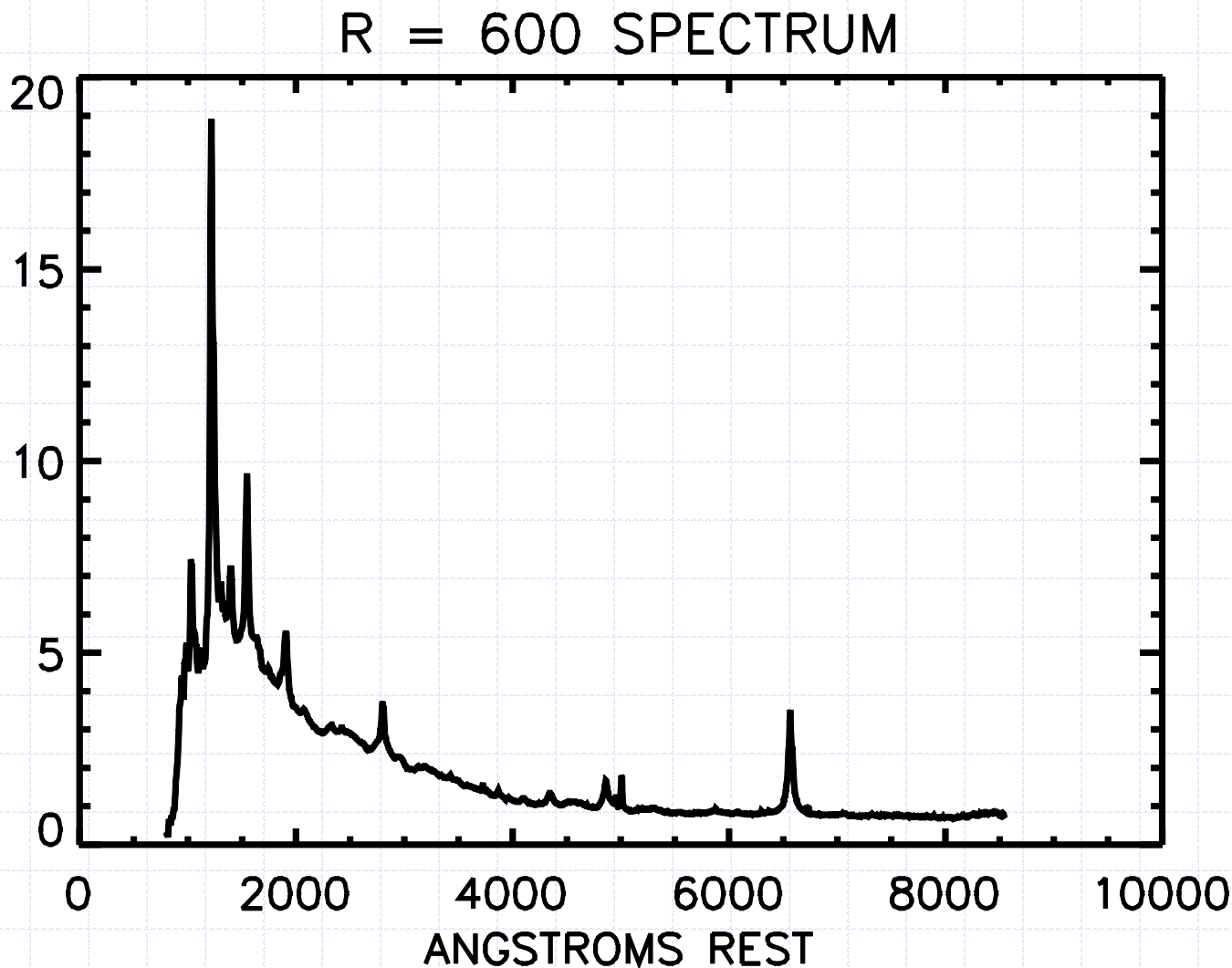


(b)

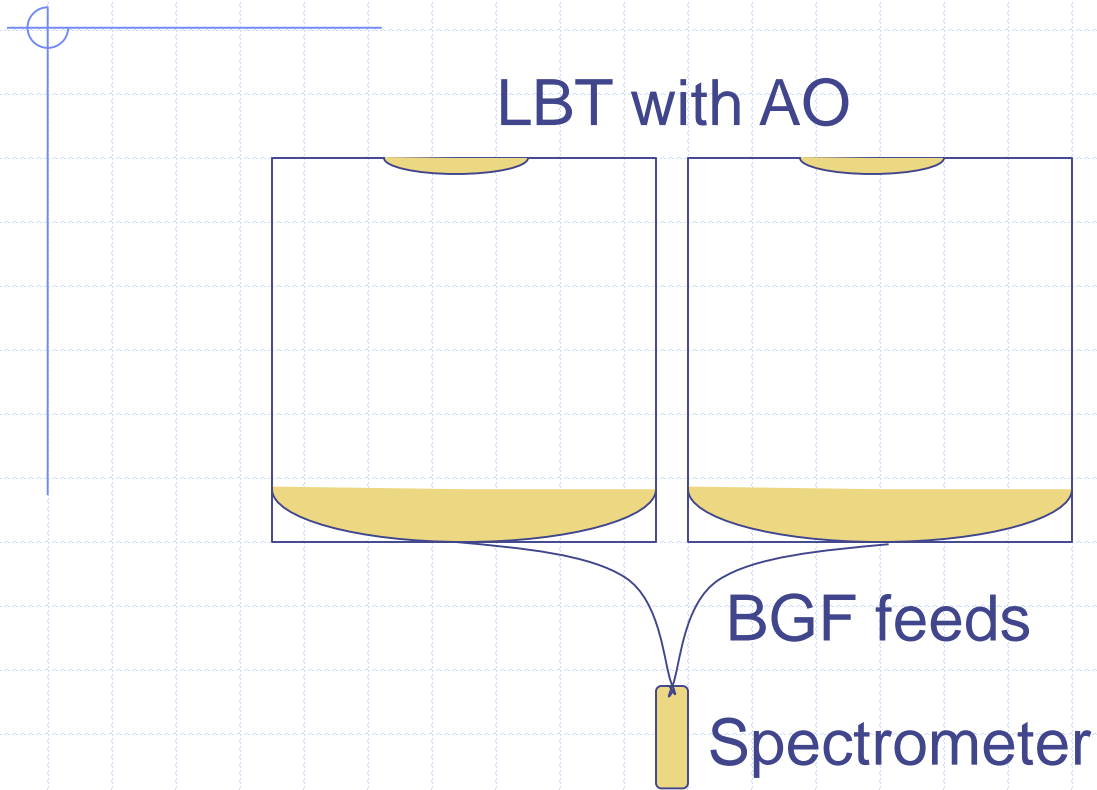
Some Science Cases

- ◆ Optical rest frame spectroscopy of QSOs
- ◆ AGN spectroscopy
- ◆ Brown Dwarf studies at greater galactic distance
- ◆ Possible exoplanet differential spectroscopy
- ◆ Spectroscopy of distant small solar system objects

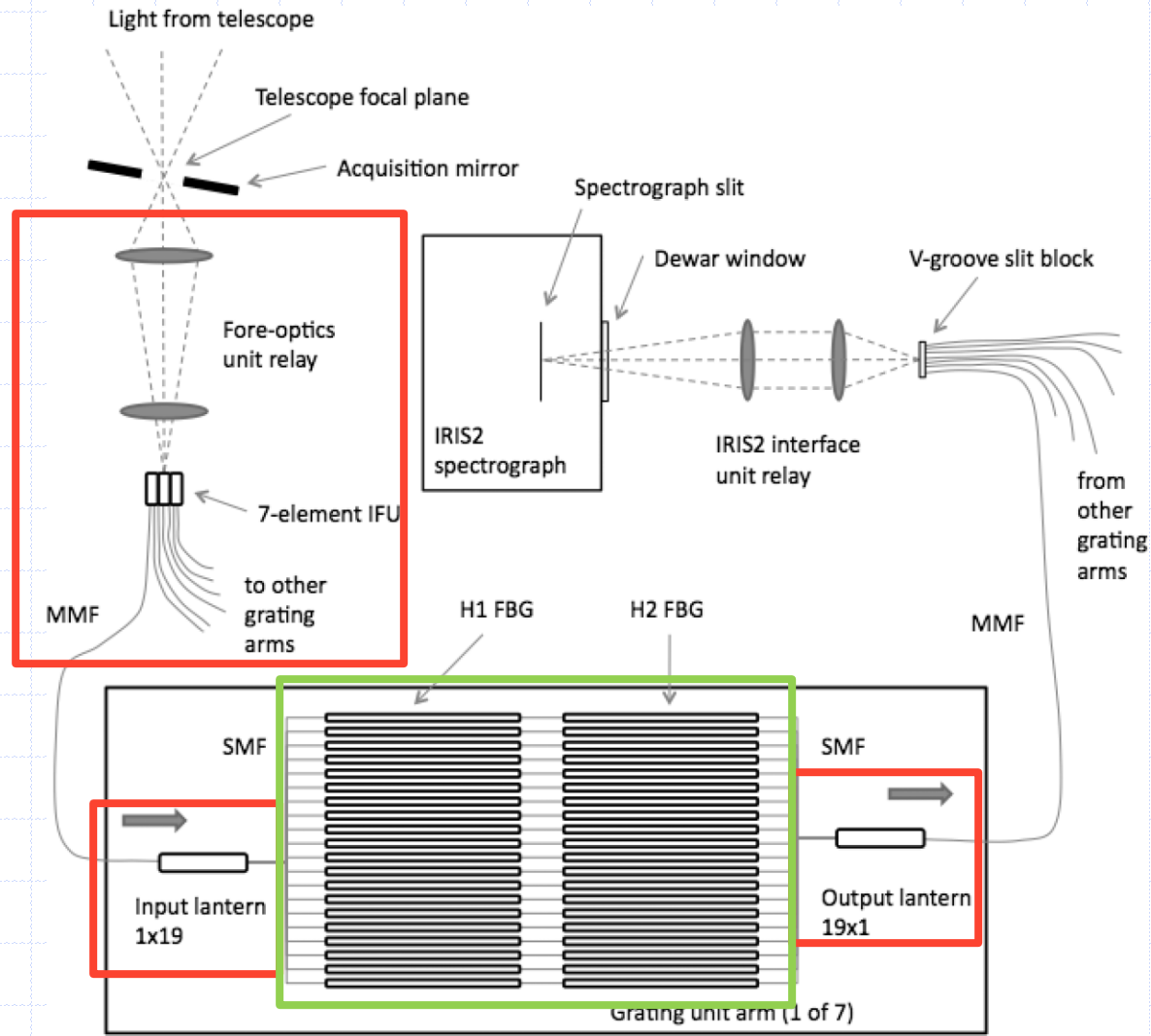
Template QSO Spectrum at R=600 @ 3000 Ang. rest



Schematic Concept



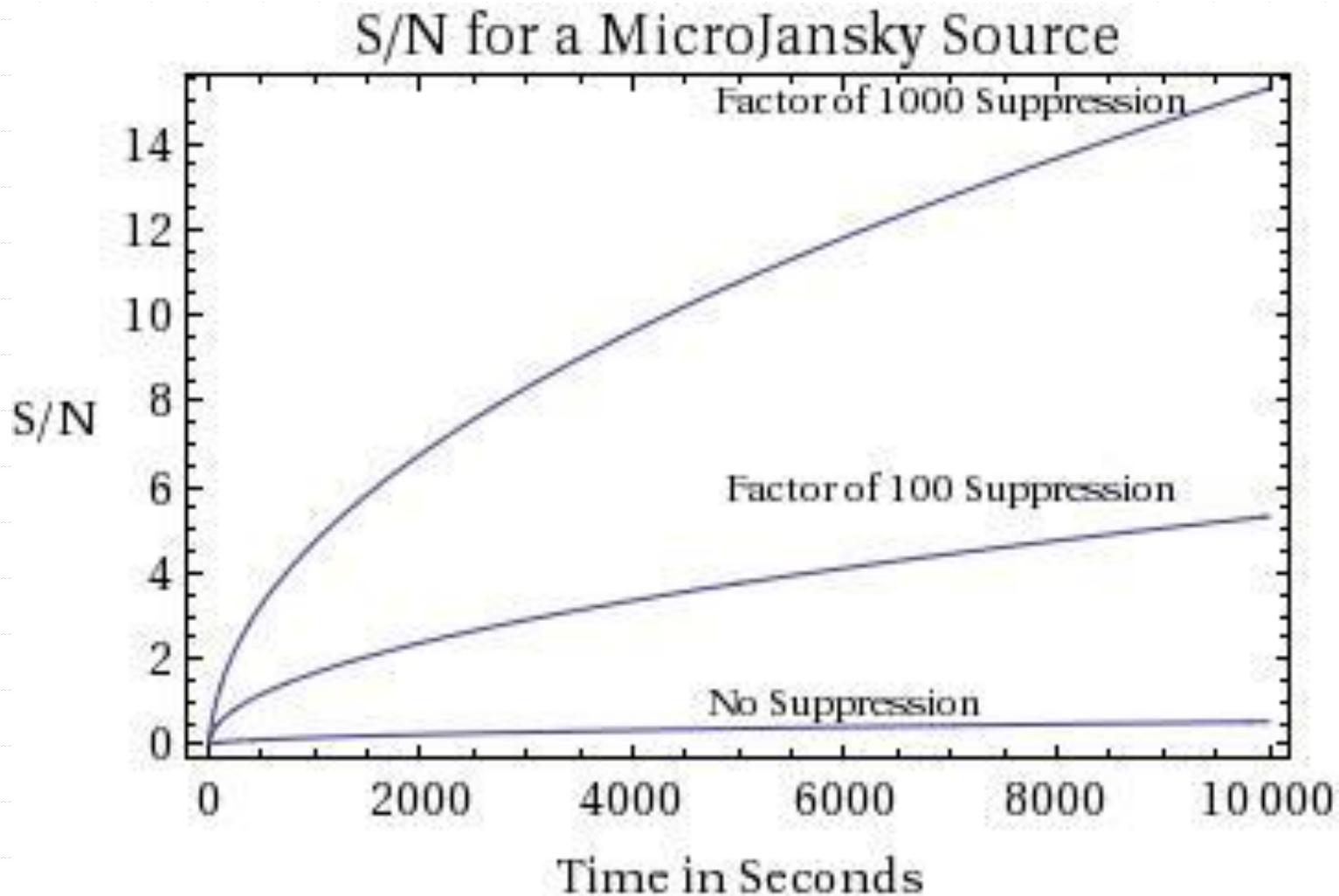
GNOSIS Spectrometer (Bland-Hawthorn et al.)



Not Needed

Two Needed

Performance (23.9 H_{AB} Mag.)



Implementation with new instrument

- ◆ Use existing 21 gr./mm grating
 - 5.5° blaze angle
 - 5" x 6" ruled area
- ◆ Use HR2G detector from JWST
- ◆ Custom collimators and camera
- ◆ 2 feed fibers with OH suppression sections
- ◆ Detachable fiber mount at LBTI focus?
- ◆ Spectrometer size $\sim 2' \times 2' \times 4'$

Additional Implementation Components

- ◆ Leach Controller
- ◆ Data acquisition computer
- ◆ Interface to LBT
- ◆ Quick look data display
- ◆ Data reduction pipeline
- ◆ Mechanical interface
 - detachable

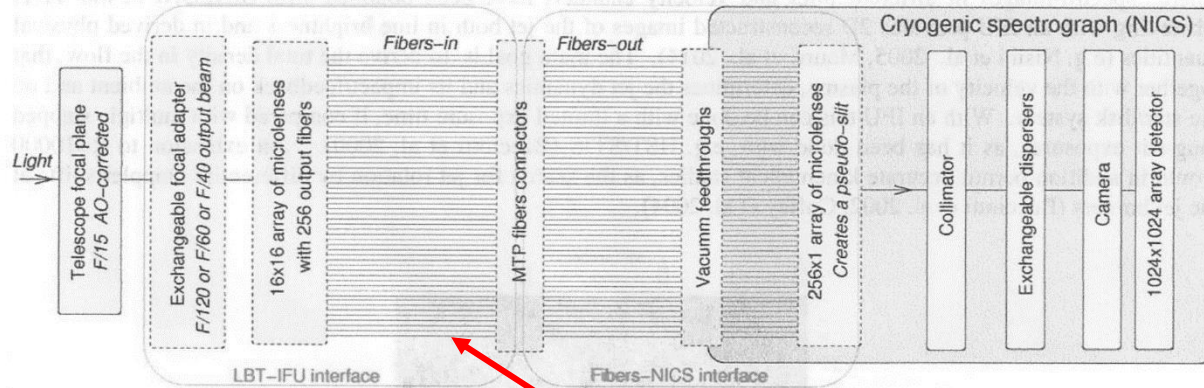
Other Implementation Opportunities

◆ iFUN@LBT

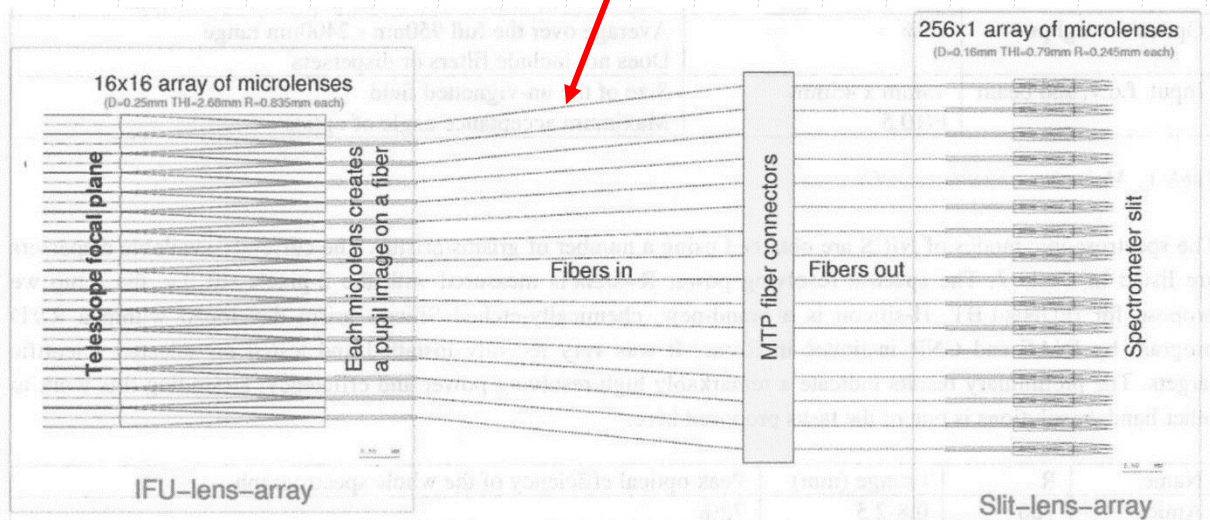
- Integral Field Unit with NICS at LBT
- Utilizes and existing spectrometer being decommissioned from TNG
- Proposed instrument for LBT led by INAF
- Multi-fiber (256) fed near infrared spectrometer

◆ Rule an OH suppression section on one of the fibers

Implementation in NICS



Replace one fiber with a BGF



Advantages

- ◆ Enables a new scientific capability not matched at any other telescope
- ◆ Low Cost
- ◆ Small size
- ◆ Can be expanded to be an integral field instrument

Uncertainties

- ◆ Previous experience (GNOSIS Bland-Hawthorn) had unexplained high background.
 - Some information that is was spectrometer rather than fiber related.
- ◆ BGF source
 - AIP has a laboratory under development and is interested in participating,
 - Previous industry source no longer interested in low volume sales.

Instrument Status

- ◆ Instrument development on hold pending more information on BGFs, their performance and availability.
- ◆ Many components, other than BGFs, already in hand.
- ◆ Since the instrument is low cost and not permanently installed it should be able to start at any time.
- ◆ A unique and powerful concept but not quite ready for prime time.